

## CHAPTER 00: EXPLORING THE HEAVENS

## NOTES AND SKETCHES

## 0.1: THE OBVIOUS VIEW

**Our Place in Space**

- ◆ Astronomy: study of the universe
- ◆ Earth does not occupy a central or special place

**Constellations In The Sky**

- ◆ About 3000 stars visible to naked eye from any random location on earth
- ◆ Constellation: pattern identified, "picture" formed by group of stars
- ◆ No astronomical significance because stars in specific constellation have no relationship to each other
- ◆ Cultural significance: important way for non-literate societies to preserve and pass on important information

**The Celestial Sphere**

- ◆ Method for locating objects in the sky
- ◆ Ignores the fact that objects are at different distances from the earth
- ◆ Project an invisible sphere out from the earth
- ◆ NCP: North Celestial Pole = projection of earth's North pole into the sky
- ◆ SCP: South Celestial Pole = projection of Earth's South pole into sky
- ◆ CE: Celestial Equator = projection of Earth's equator into the sky
- ◆ Stars appear fixed with respect to each other, but entire celestial sphere "spins" as Earth turns on its axis

**Celestial Coordinates**

- ◆ To precisely locate an object, two coordinates required
- ◆ On Earth: latitude measures degrees of angle from  $0^\circ$  to  $90^\circ$  north or south of equator
- ◆ In space: declination measured in degrees from  $0^\circ$  to  $90^\circ$  north (+) or south (-) of celestial equator
- ◆ On Earth: longitude measured in degrees around the equator
- ◆ Zero chosen arbitrarily:  $0^\circ$  longitude = Royal Observatory, Greenwich, England
- ◆ Measure from  $0^\circ$  to  $180^\circ$  East (towards Asia) or  $0^\circ$  to  $180^\circ$  West (towards N America)
- ◆ In space: right ascension measured in hours, minutes, seconds around the celestial equator
- ◆ Zero chosen arbitrarily: 0h RA = where sun crosses the celestial equator on vernal equinox (HUH?!?!?)
- ◆ Measure from 0h to 24 h RA around CE in the same direction as the earth spins

## 0.2: EARTH'S ORBITAL MOTION

**Day-to-Day Changes**

- ◆ Solar day: Earth completes one rotation with respect to the sun = 24 hours
- ◆ Start timing when sun crosses S meridian (noon), stop timing when sun crosses meridian again tomorrow
- ◆ Sidereal day: Earth completes one rotation with respect to distant star = 23h56m
- ◆ Start timing when star crosses S meridian, stop timing when same star crosses meridian again tomorrow

## NOTES AND SKETCHES

**Seasonal Changes**

- ◆ Earth's axis is tilted
- ◆ Draw an enormous flat plane through the equator of the sun, and extend it all the way through the solar system
- ◆ This plane does not cut the Earth in half at its equator
- ◆ Axis tilt =  $23.5^\circ$

**Ecliptic**

- ◆ Apparent path of the sun across the sky (really the Earth in motion, not the sun)
- ◆ Ecliptic is tilted b/c Earth's axis is tilted

**Seasons**

- ◆ Tilt of axis creates changing seasons
- ◆ Distance from sun does not create seasonal changes
- ◆ Earth is marginally closer to sun in Dec than Jun, but Dec is not the hottest month in the N hemisphere!

**Summer Solstice**

- ◆ Usually 06/21 (may fall  $\pm$  a calendar day)
- ◆ Longest day of the year (N hemisphere) because sun has maximum declination ( $+23.5^\circ$ )
- ◆ N hemisphere is tipped toward sun, more direct daylight makes the season summer

**Winter Solstice**

- ◆ Usually 12/21 (may fall  $\pm$  a calendar day)
- ◆ Shortest day of the year (N hemisphere) because sun has minimum declination ( $-23.5^\circ$ )
- ◆ N hemisphere is tipped away from sun, less direct daylight makes season winter

**Equinoxes**

- ◆ Vernal (Spring) Equinox: 03/21 (may fall  $\pm$  a calendar day)
- ◆ Autumnal Equinox: 09/21 (may fall  $\pm$  a calendar day)
- ◆ Equal length day & night because sun crosses CE (dec =  $0^\circ$ )

**Long-Term Changes**

- ◆ Earth's axis wobbles slightly as it spins
- ◆ Today, NCP points almost perfectly at Polaris
- ◆ Wobble means that Polaris was not always the Pole Star, and will not be forever

**0.3: THE MOTION OF THE MOON****Lunar Phases**

- ◆ Understanding the phases helps us really start to get to grips with the layout of the solar system
- ◆ New moon: moon located in between Earth and sun (angle is  $0^\circ$ )
- ◆ 1st quarter: moon makes a  $90^\circ$  angle (If you are the Earth, stick your right arm straight out and make a fist: your fist is the sun. Stick your left arm straight out to your side, so your arms make a  $90^\circ$  angle. Your left fist is the moon.)
- ◆ Full moon: moon is  $180^\circ$  away from sun in sky (If you are the Earth, the sun is directly in front of you, then the moon would be directly behind you.)
- ◆ 3rd quarter: moon makes  $90^\circ$  angle (switch hands, and let your left fist be the sun sticking straight out in front. Make a  $90^\circ$  with your right arm, and your right fist is the moon)

## NOTES AND SKETCHES

**Lunar Month**

- ◆ Sidereal Month: 27.3 days for the moon to complete one full rotation with respect to distant star
- ◆ Synodic Month: 29.5 days for moon to complete one full cycle of phases, or a complete rotation with respect to the sun (as seen from the earth)

**Solar Eclipse**

- ◆ Sun is eclipsed by the moon: moon passes in between Earth and sun
- ◆ Can only happen when phase of moon is new
- ◆ Does not happen every month because moon's orbit is tilted with respect to ecliptic
- ◆ Annular eclipse: moon is farthest from Earth, making it appear slightly smaller (so it does not completely cover solar disk)

**Lunar Eclipse**

- ◆ Shadow of the Earth eclipses the moon: Earth passes in between sun and moon
- ◆ Can only happen when moon is full
- ◆ More frequent occurrence than solar eclipse
- ◆ Partial eclipses not uncommon

**0.4: THE MEASUREMENT OF DISTANCE****Triangulation**

- ◆ Measure distance to objects that are too far or inconvenient to be measured directly
- ◆ Requires some geometry and trigonometry
- ◆ Observe the same object from two different vantage points, compare
- ◆ This works for stationary objects on Earth, but can also be used to locate planets

**Parallax**

- ◆ Apparent shift in the position of an object in the foreground with respect to the background
- ◆ Result of changing point of observation, not the motion of the actual object
- ◆ For close object, large parallax observed with relatively small baseline shift
- ◆ The farther an object, smaller the parallax: increase baseline to increase parallax

**0.5: SCIENTIFIC THEORY AND SCIENTIFIC METHOD****It's Only a Theory...**

- ◆ A scientific theory has been repeatedly tested, and never found to be false (not once, not even a little bit)
- ◆ If something is referred to by scientists as a theory, it is widely accepted as the best framework for explaining something
- ◆ A theory must be able to explain what has been observed and predict what should happen as a consequence
- ◆ Scientists accept that, if new facts or experiments reveal a flaw, the theory must be modified or discarded

**Scientific Method**

- ◆ Process by which science gets done
- ◆ The whole point is, it's never actually done

**Sizing Up Planet Earth**

- ◆ Eratosthenes accurately measured Earth's circumference and diameter in about 200BC
- ◆ All you need are two sticks and scratch paper
- ◆ More to the point: He apparently took it for granted that the Earth was a sphere...in 200BC